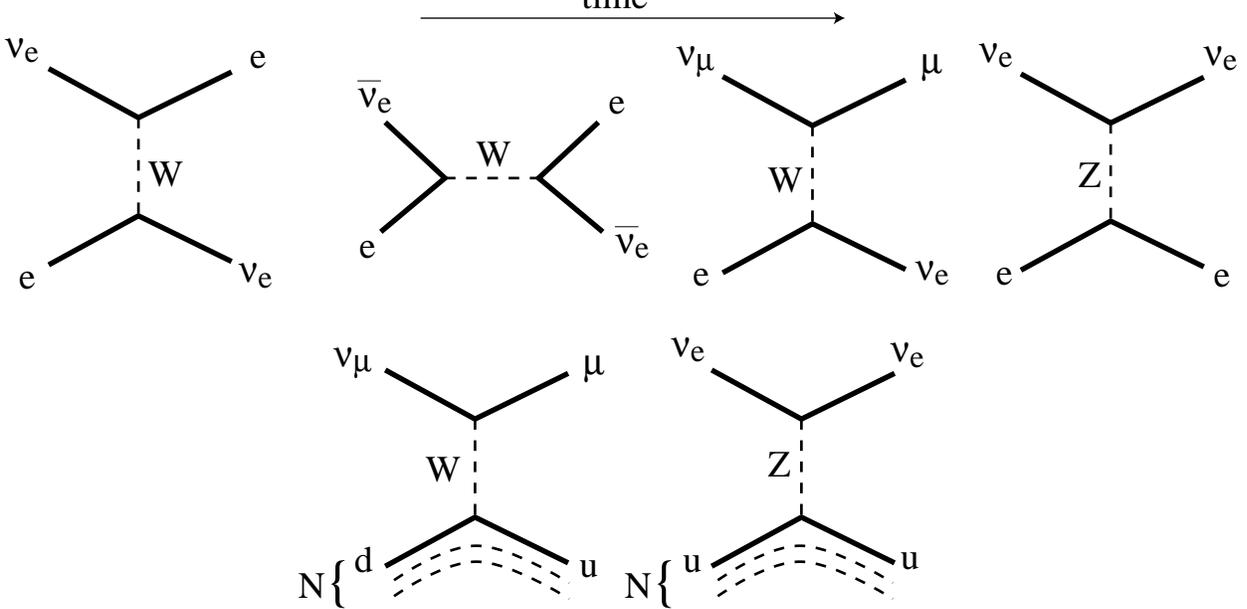


# Neutrino shower physics

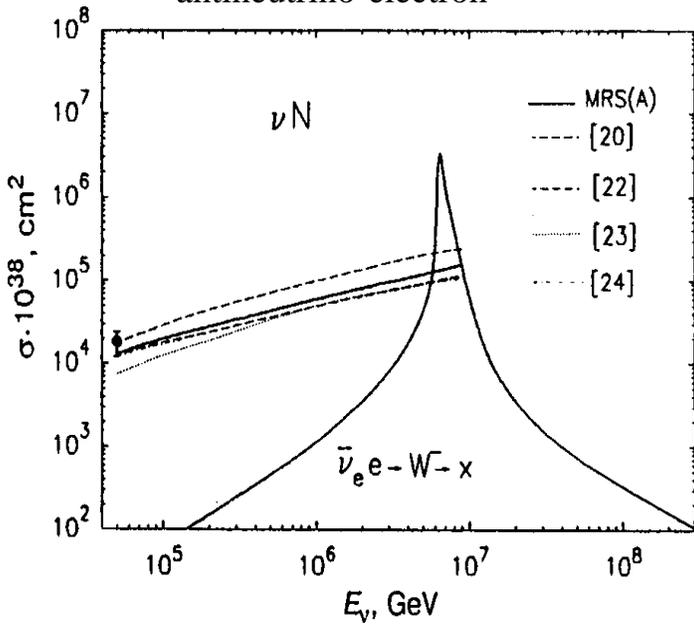
Shower consists of the following components:

- hadronic (pions, kaons etc.)
- electromagnetic (electrons, positrons, gamma photons)
- muons

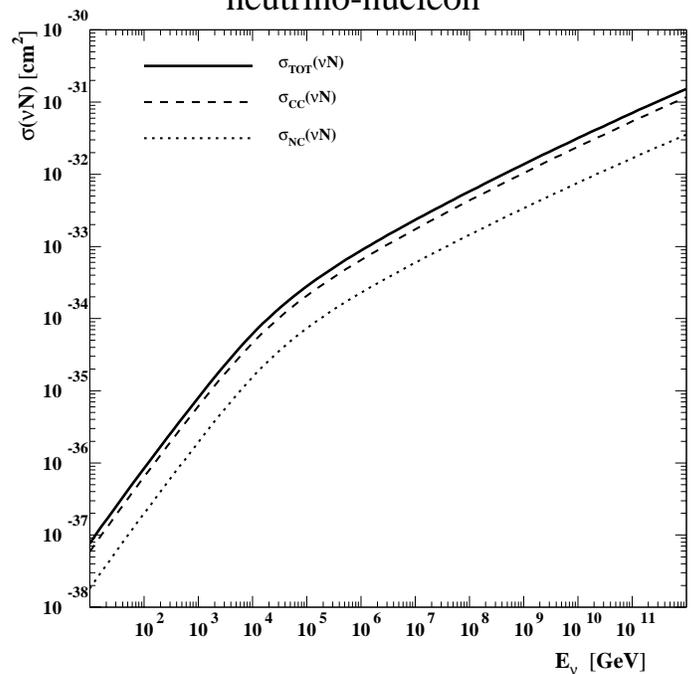
## Neutrino cross-section



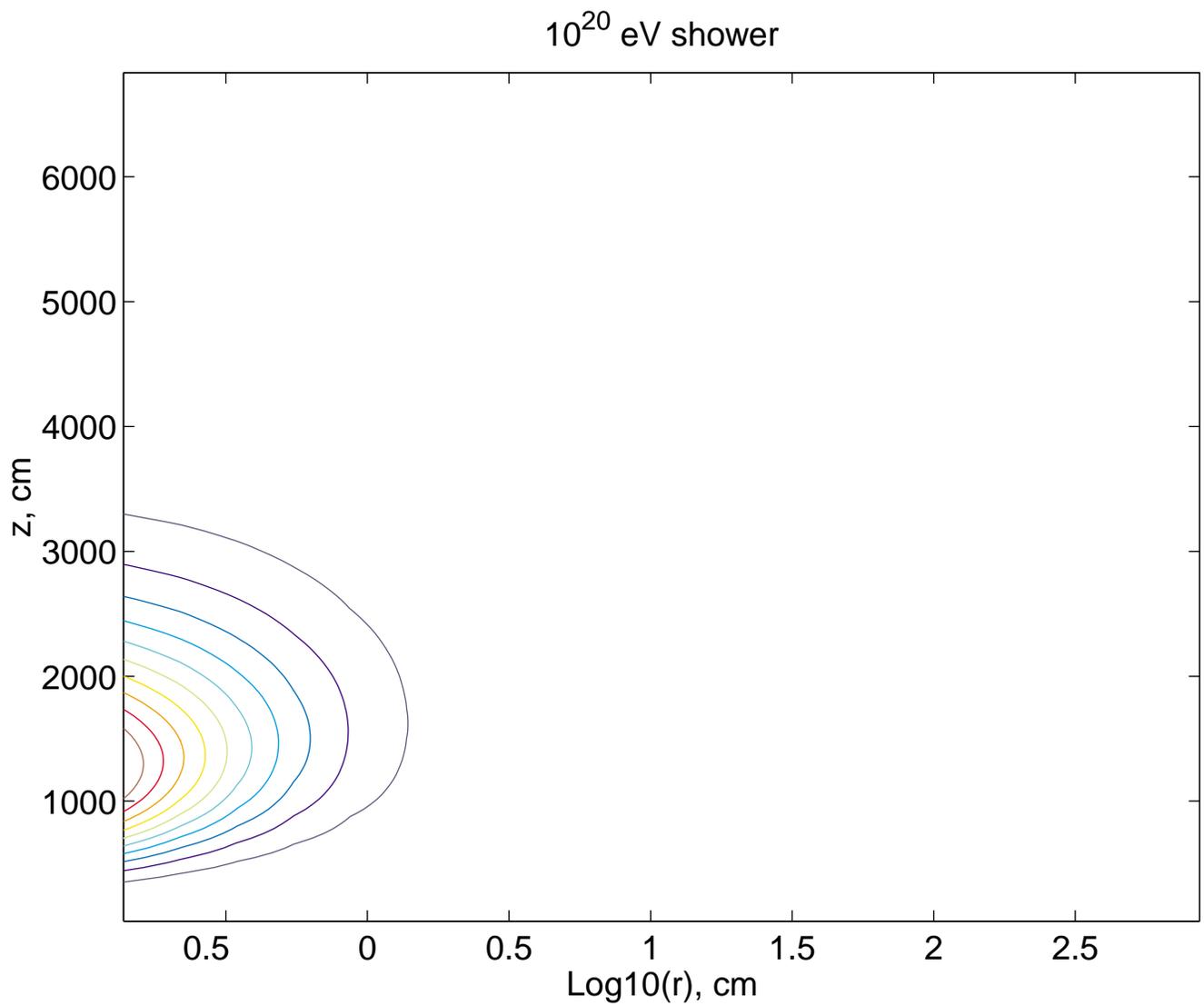
Butkevich et al., 1998:  
neutrino/antineutrino-nucleon  
antineutrino-electron



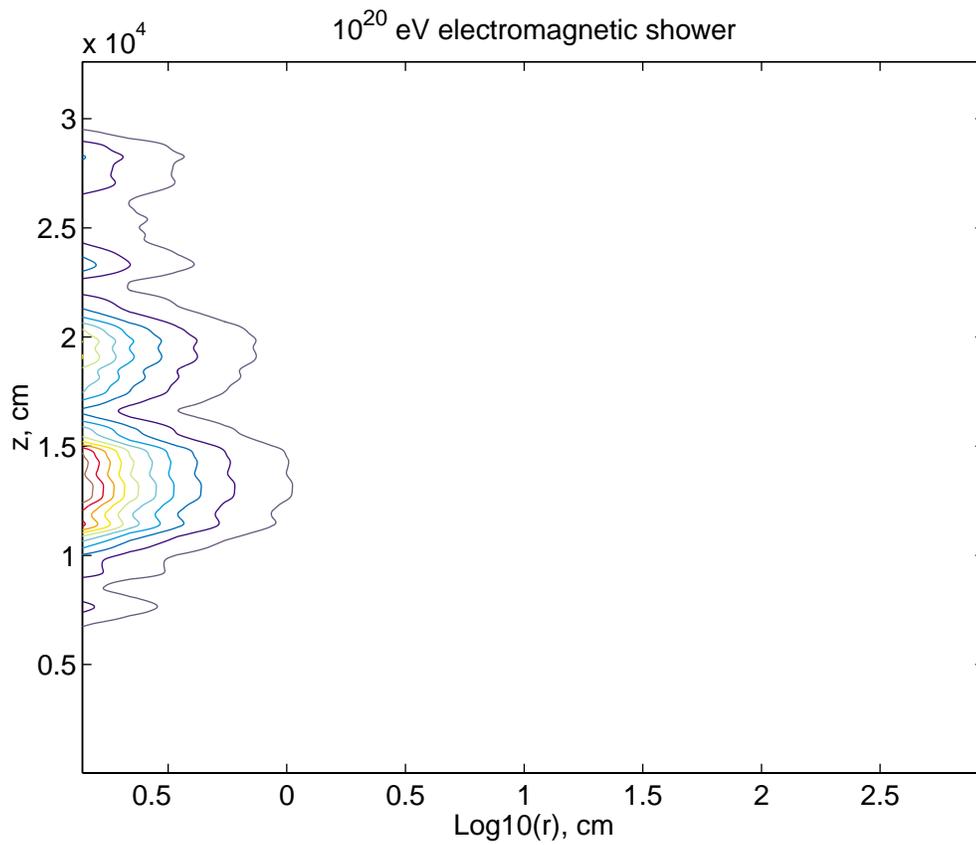
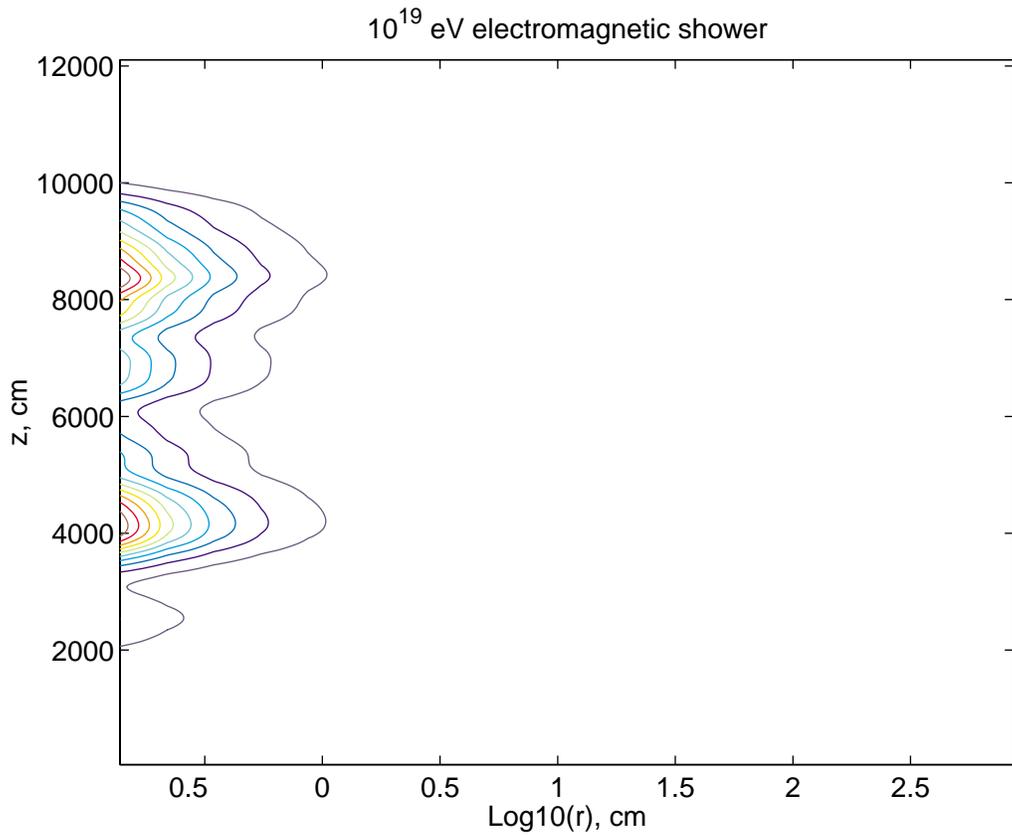
Kwiecinski et al., 1998:  
neutrino-nucleon



# Analytical shower parametrization [Learned, 1979]

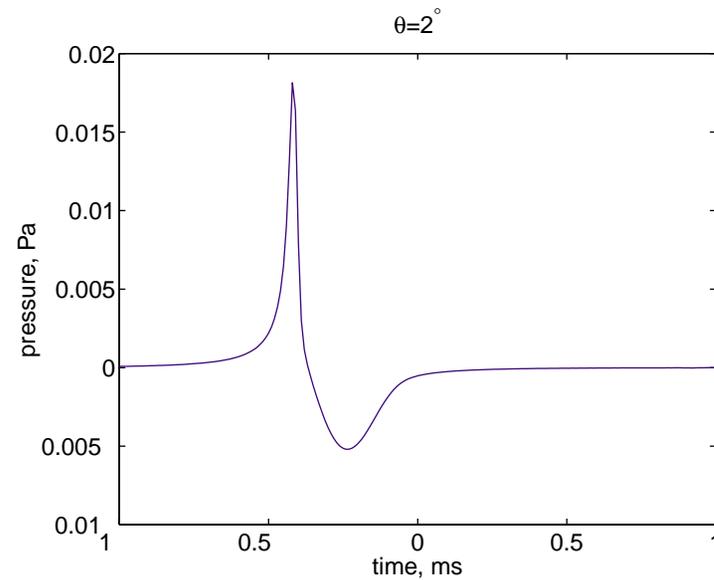
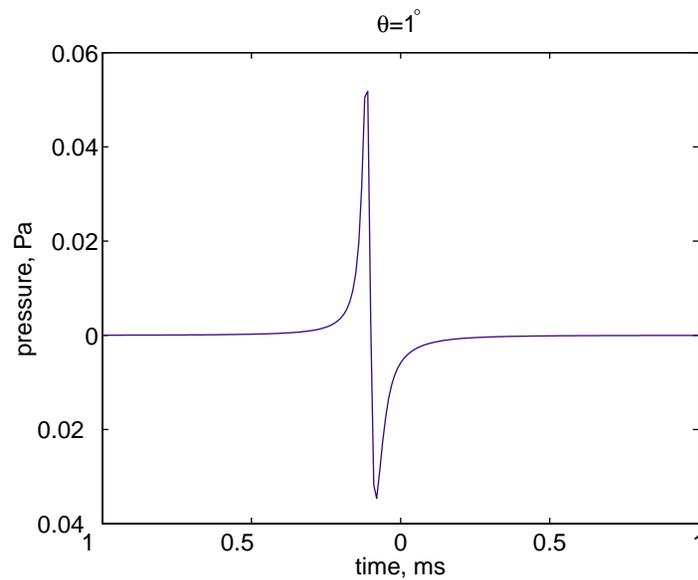
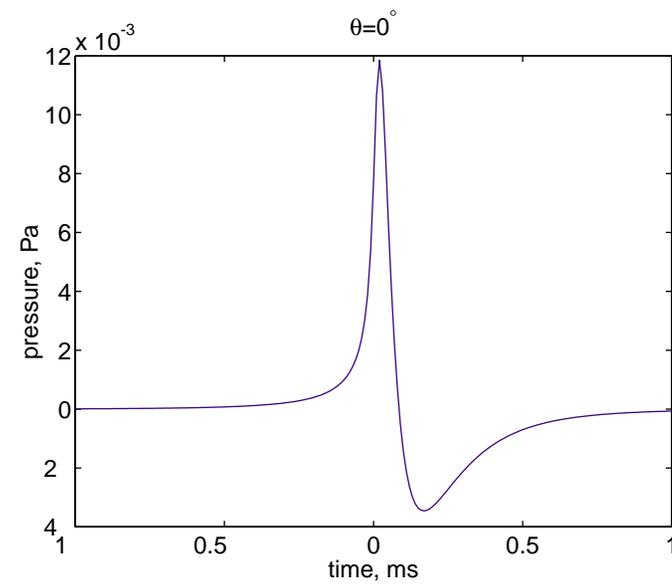
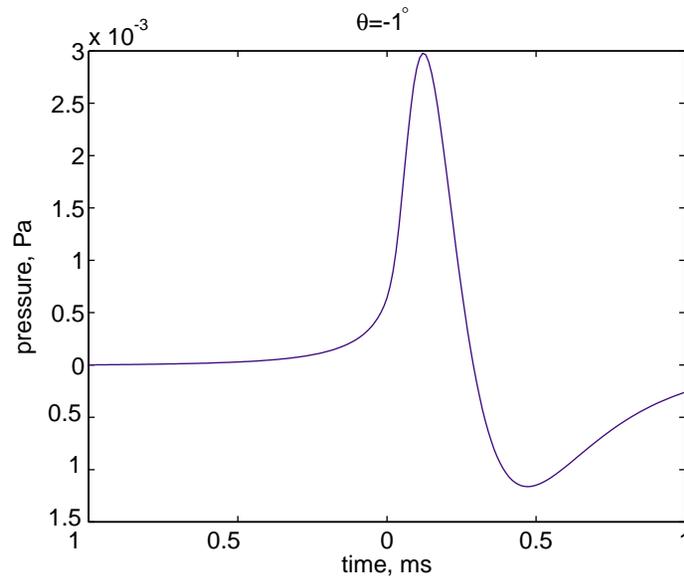
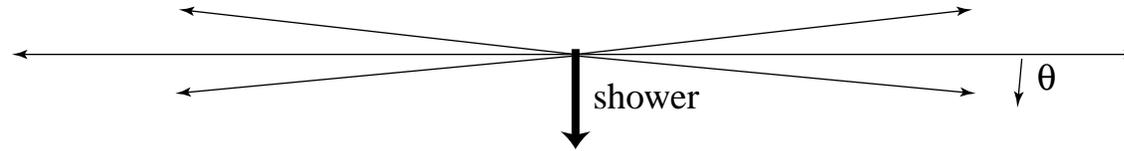


# Alvarez-Muniz et al. model results

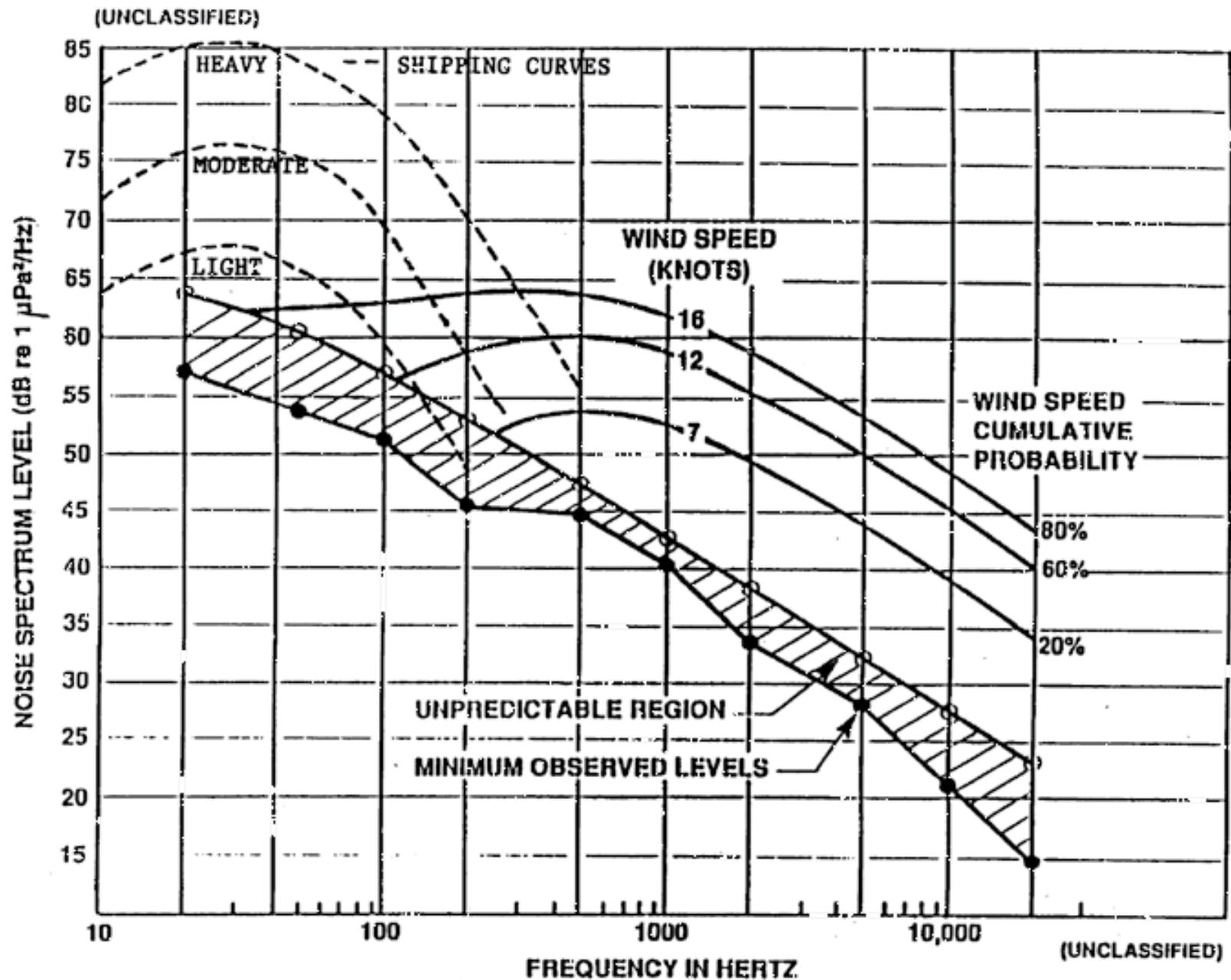


# Sound pulse simulation results (for Learned parametrization)

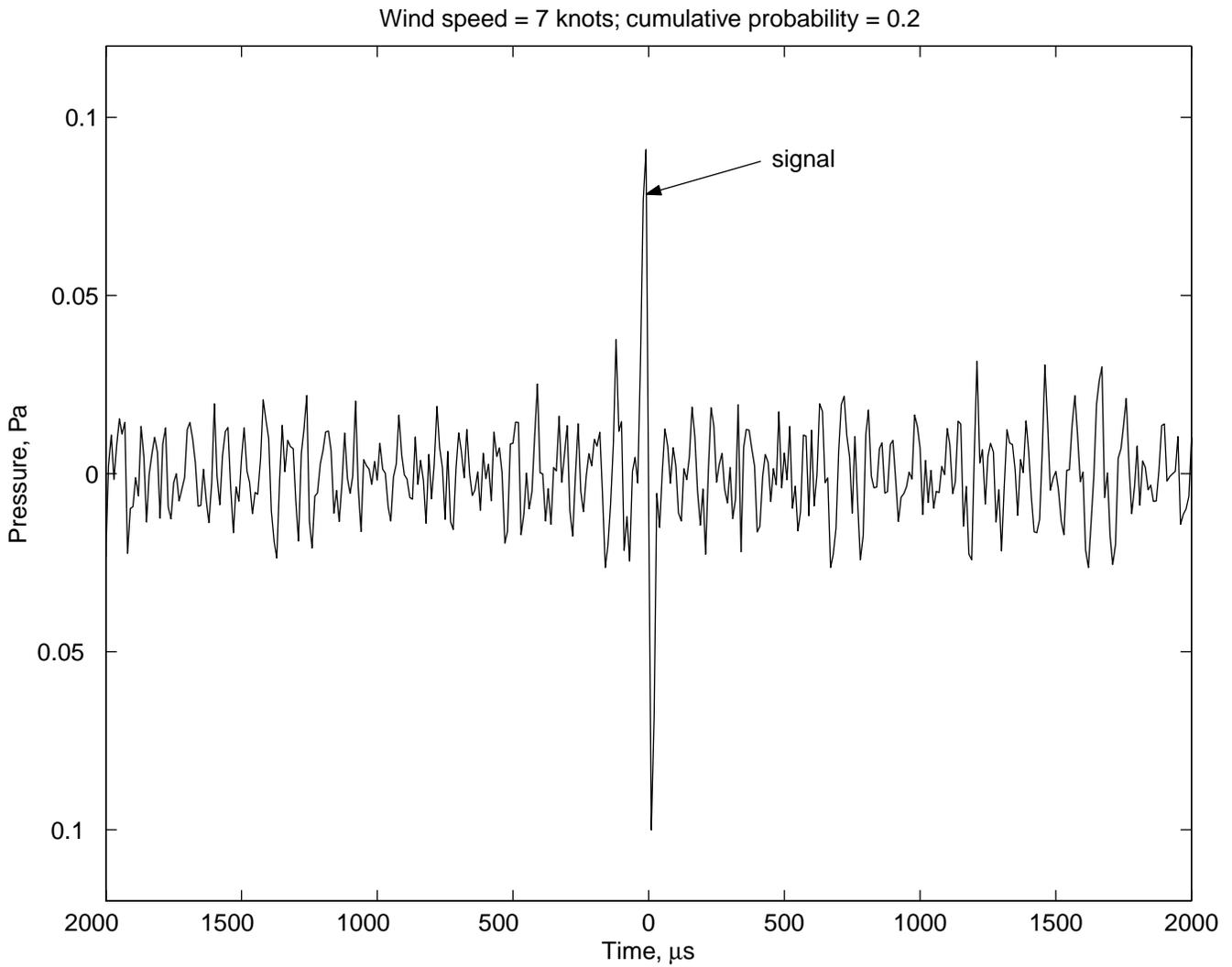
The acoustic radiation is in a narrow disk perpendicular to shower axis



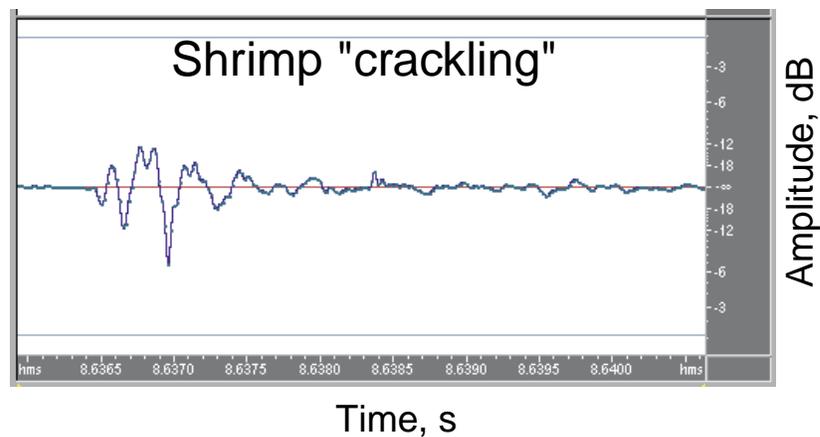
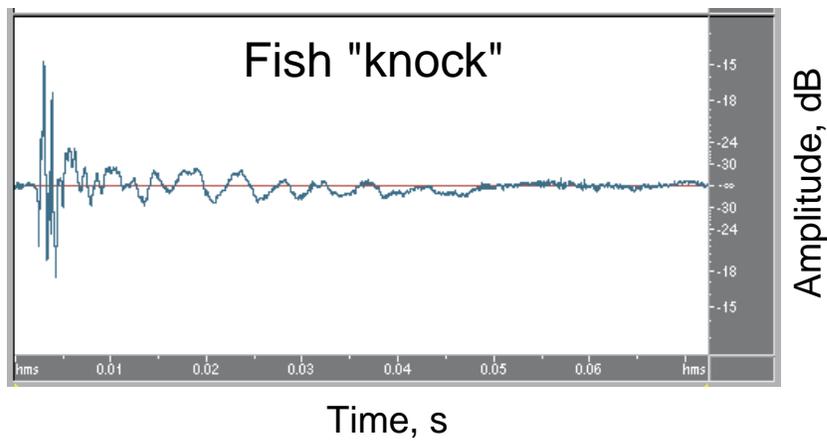
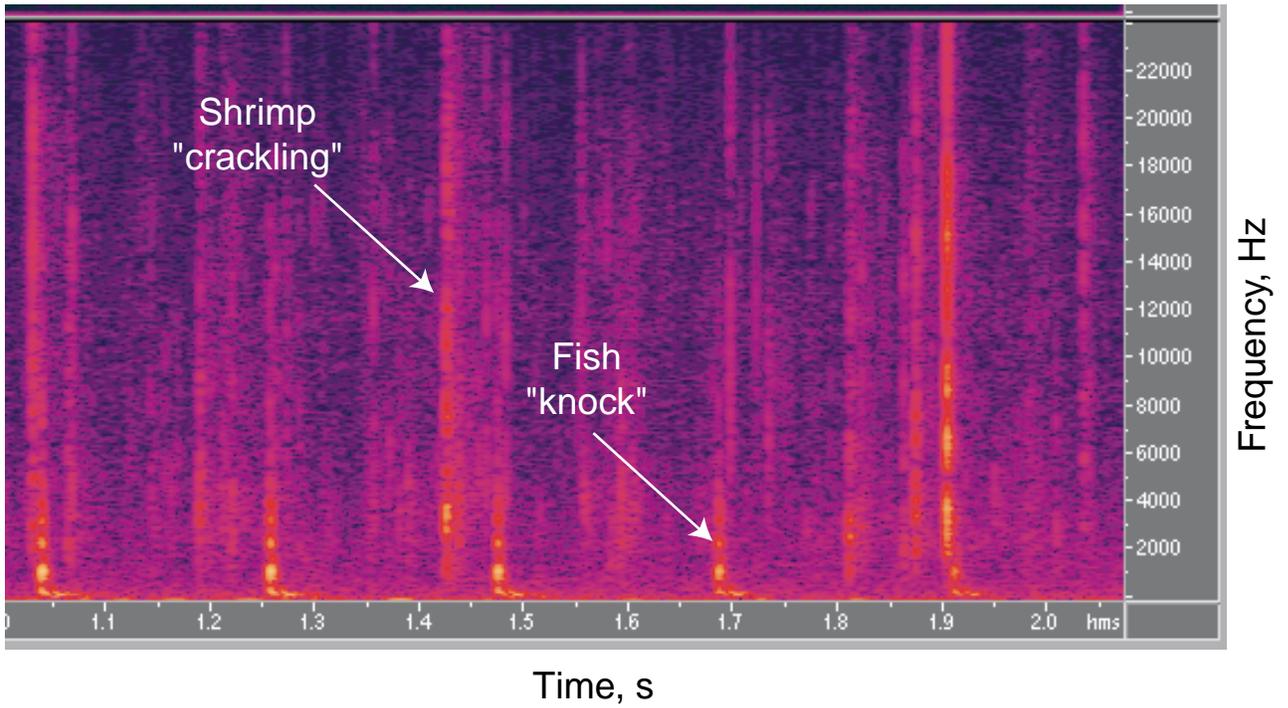
# AMBIENT NOISE LEVELS AT AUTEC, TOTO, BAHAMAS



# The pulse of $10^{20}$ eV neutrino recorded at 1 km distance super-imposed on the statistical noise

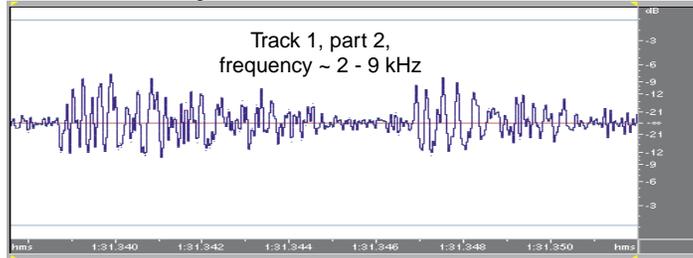


# Examples of animal noise (from UCSD tape)



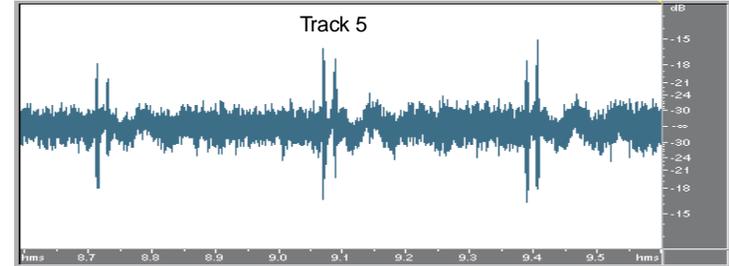
# More examples of animal noise (from AUTECD CD)

### Sperm whale "click"



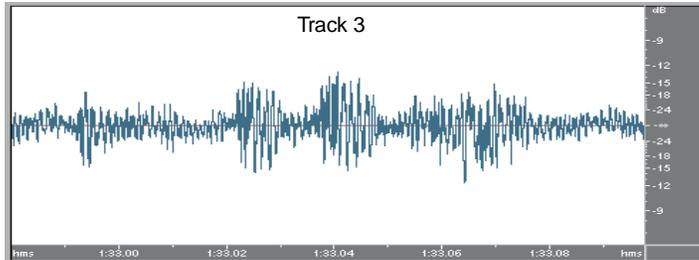
Amplitude, dB

### Pilot whale "knocks"

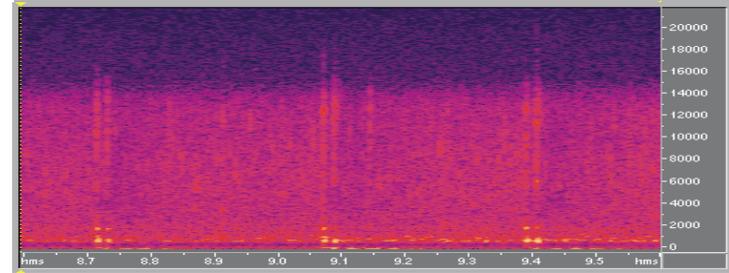


Amplitude, dB

### Pilot whale "rattle"

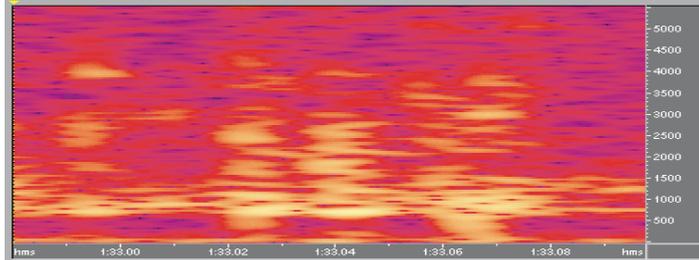


Amplitude, dB



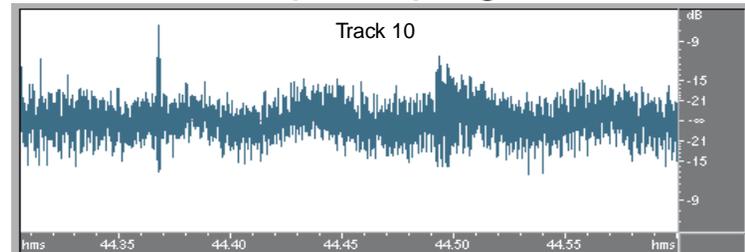
Frequency, Hz

Time, s



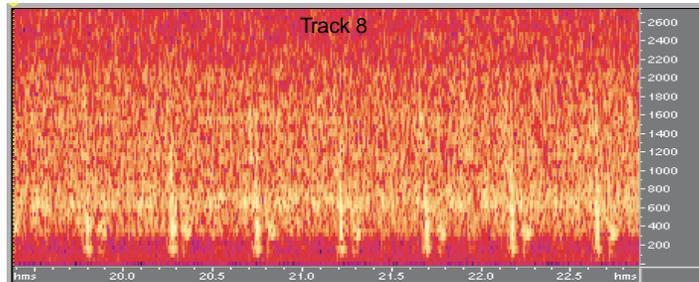
Frequency, Hz

### Dolphin "pinger"

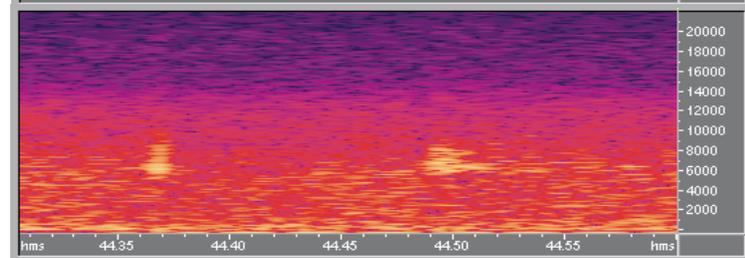


Amplitude, dB

### Minke whale "thump"



Frequency, Hz



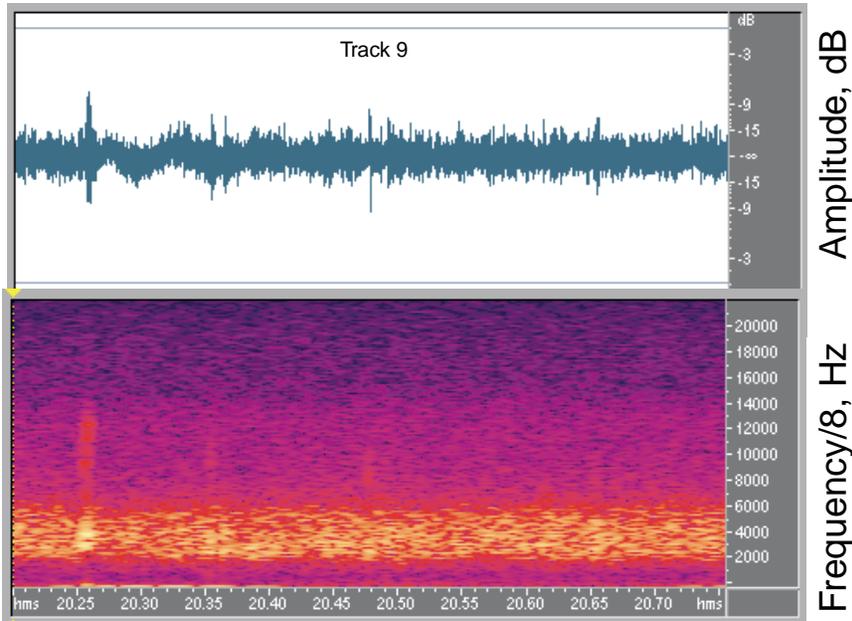
Frequency/4, Hz

Time, s

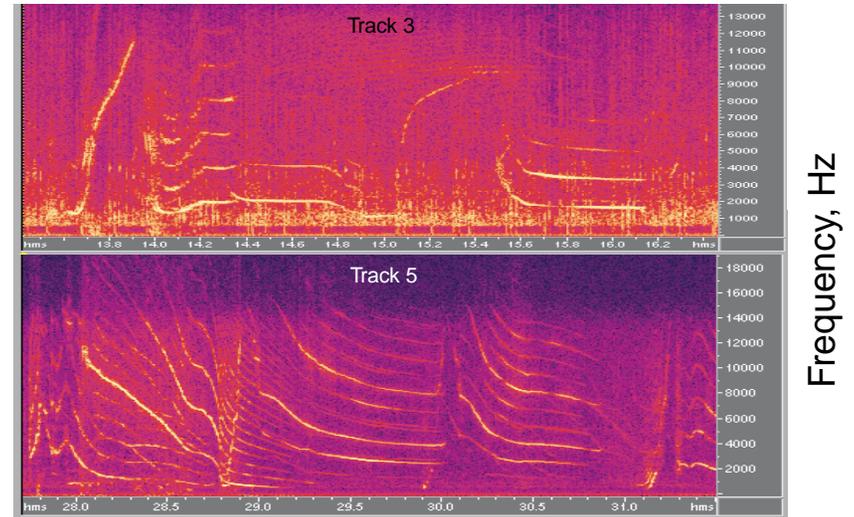
Time x 4, s

# Still more examples from AUTECD CD

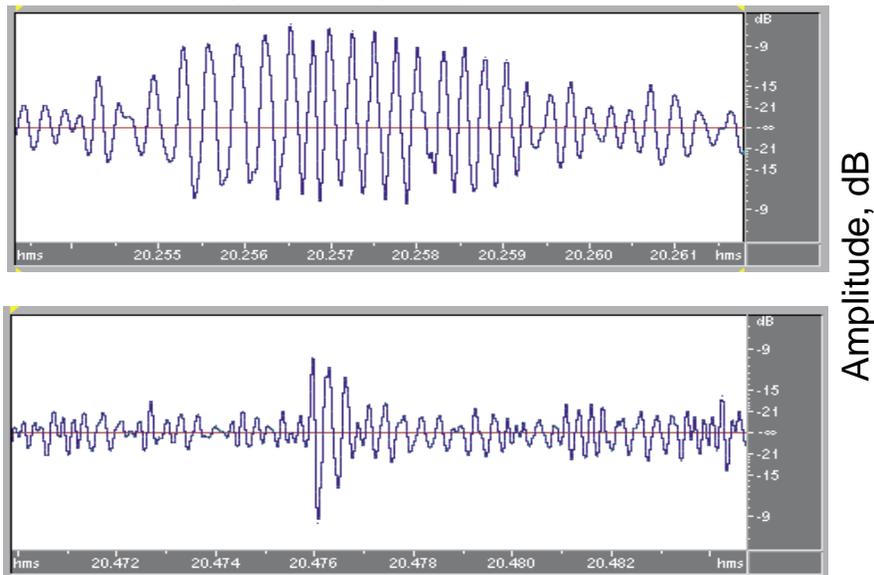
## Whale "pinger"



## Non-pulsed sounds: Pilot whale "whistle"

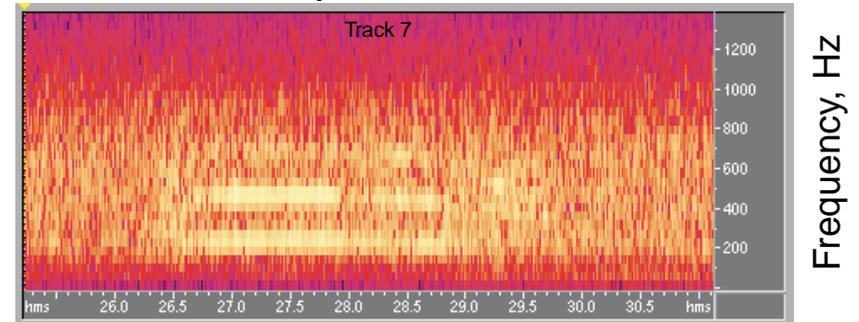


## Separate pinger trains

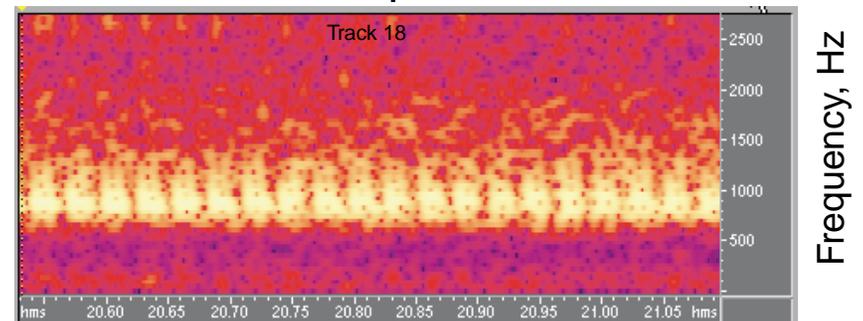


Time x 8, s

## Humpback whale



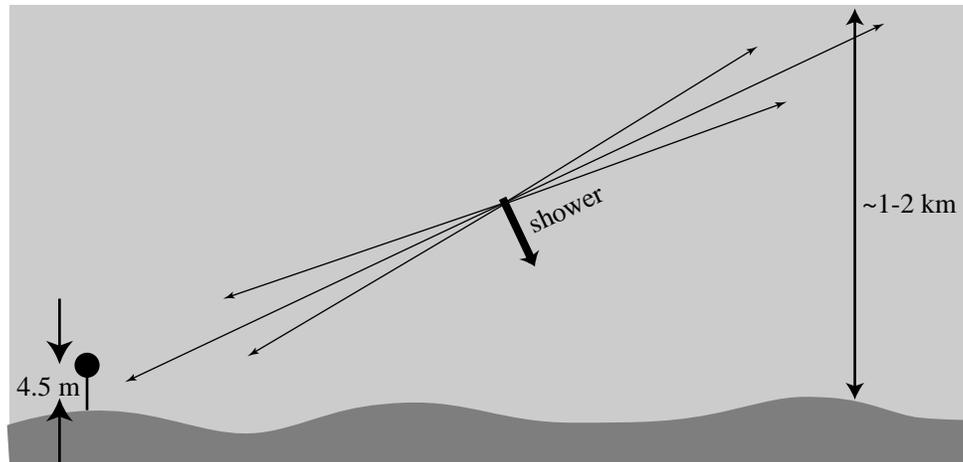
## Squid



Time, s

# Problems with acoustic detection of neutrinos to be solved

1. Narrowness of the detection volume => hard to correlate between hydrophones.



2. Ambient Gaussian noise =>

- only detection of the highest energies neutrinos?
- use of a sophisticated detection technique (wavelet analysis).

Note that another analogous proposed experiment, namely SADKO [Sea Acoustic Detector of Cosmic Objects, Butkevich et al., 1992], to be located in Mediterranean at 4 km deep, has ~10 times smaller amplitude of ambient Gaussian noise [Dedenko et al., 1994].

3. The variable temperature profile => the amplitude depends on the location of the shower. Possible use of the nonthermal acoustic emission mechanisms.

(temperature profiles)

4. How to infer the location of the shower:

The direction can be guessed using both direct signal and the signal reflected from the bottom (the hydrophone is located at ~4.5 m from the bottom).